

# Energie noire et grille de calcul Le projet ZEN

*A. Tilquin*

*Centre de Physique des Particules de Marseille*

[www.eu-egee.org](http://www.eu-egee.org)



- **Scientific Goal:**

Address open questions of fundamental cosmology:  
dark matter/dark energy sector and primordial universe  
Construct a public tool....

- **Main people involve:**

ANR program

CPPM

Alain Bonnissent, Anne Ealet, Dominique Fouchez, Lei Sun, Diane Talon-Esmieu, Charling.  
Tao

CEA

Philippe.Brax, Jean Batiste Melin, Christophe.Yèche, Dominique.Yvon, Nathalie Palanque-Delabrouille, Alexandre.Réfrégier, Jim.Rich

CPT

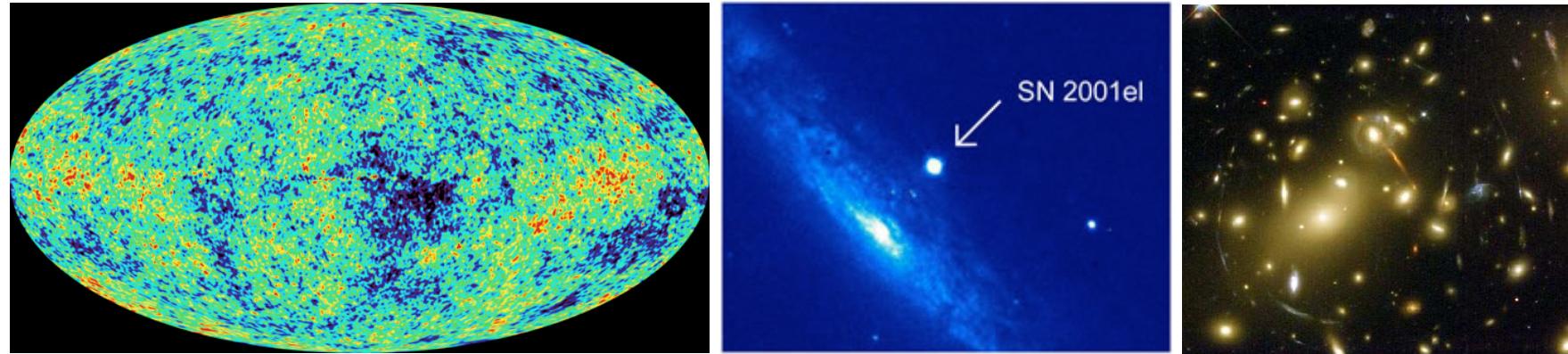
Christian. Marinoni, Pierre Taxil, Jean-Marc Virey, Stefan. Linden  
LAL

Francois Couchot, Olivier Perdereau, L. Perotto, S. Plaszczynski , C. Rosset

FCPPL: Franco-China-Particle-Physics-Laboratory program

IHEP/PKU/NAOC/Tsinghua U/Beijing Normal U.

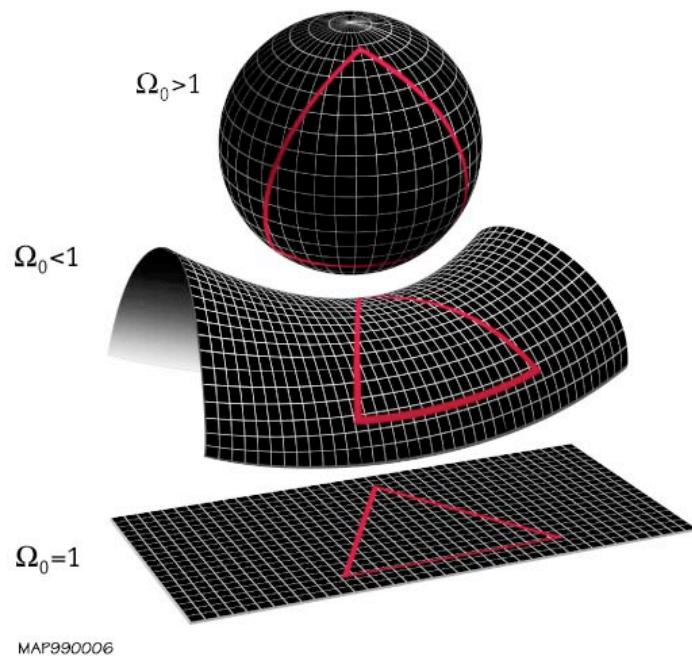
ZHANG XinMin, ZHAO Gongbo , QIAN Zuxuan , XIA Junqing, FAN Zuhui, LI Hong , Zhu  
Zhonghong, QIN Bo, DENG Jinsong , CHAO Wu , ZHOU Xu , WANG Xiaofeng



- Few words about cosmology and dark energy
- How to characterize dark energy
- Statistical method and datagrid
- Results already obtained by the Chinese and French groups (ESR and Euchina)
- Prospectives

- Cosmology based on:  
Homogenous and isotropic Universe

General relativity:  $G_{\mu\nu} = 8\pi G T_{\mu\nu}$



Energy content:  $\Omega = \rho/\rho_c$

Equation of state:  $w(z) = p/\rho$

Matter ( $\Omega_m$ ) :  $w = 0$

Radiation ( $\Omega_r$ ) :  $w = 1/3$

Cosmological cste ( $\Omega_\Lambda$ ):  $w = -1$

Dark energy ( $\Omega_X$ ) :  $w(z)$

$$\Omega_T = \Omega_m + \Omega_r + \Omega_X \quad (\rho_c = 10^{-29} \text{ g/cm}^3)$$

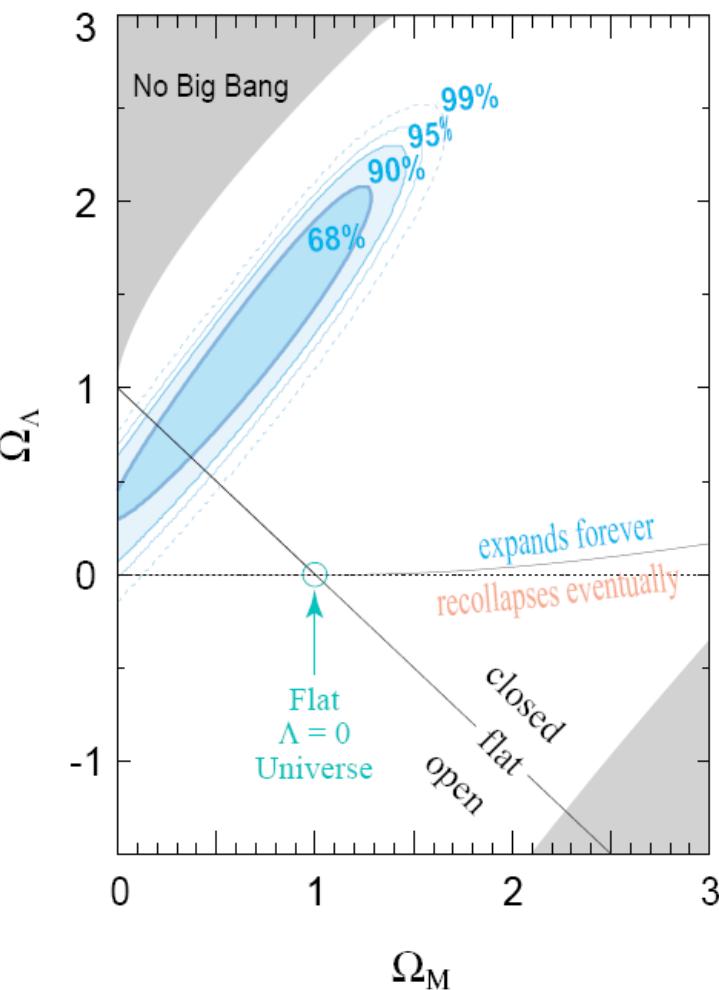
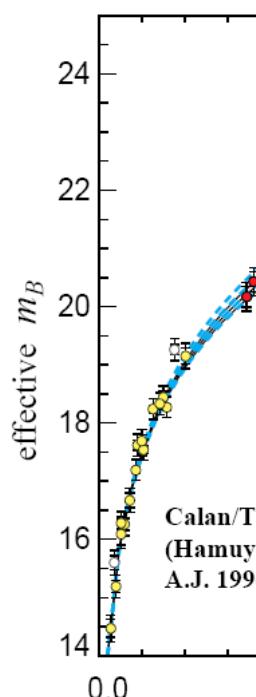
# Some words about dark energy

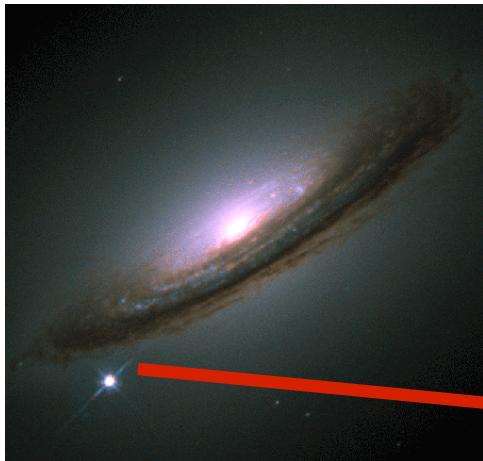
- In 1998, the Supernova Cosmology Project and High-z team shown that high red-shift supernovae are fainter than expected: a new energy component is needed.

Dark energy or cosmological constant characterized by reduce density:  $\Omega_\Lambda = \rho_\Lambda / \rho_c$

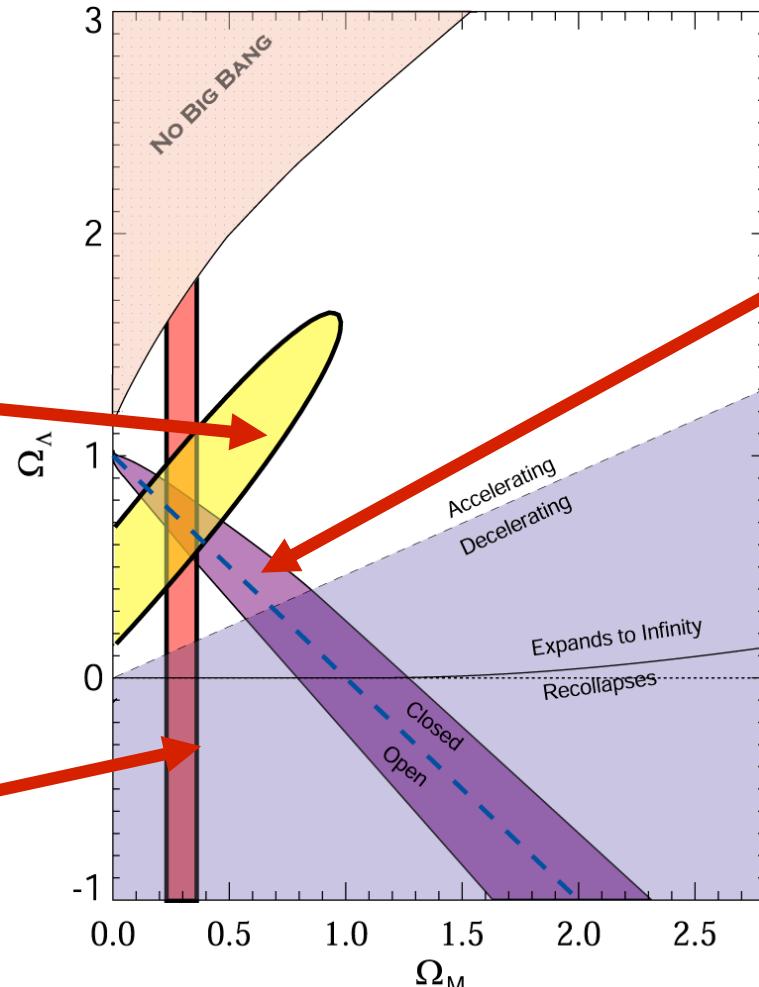
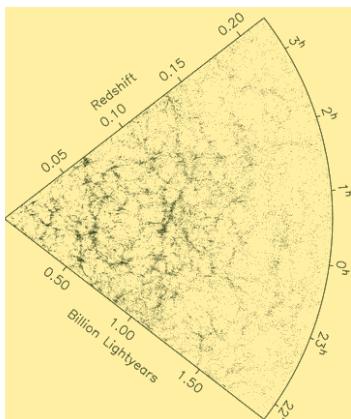
For a flat Universe:

$$\Omega_\Lambda = 0.72^{+0.09+0.05}_{-0.08-0.04}$$



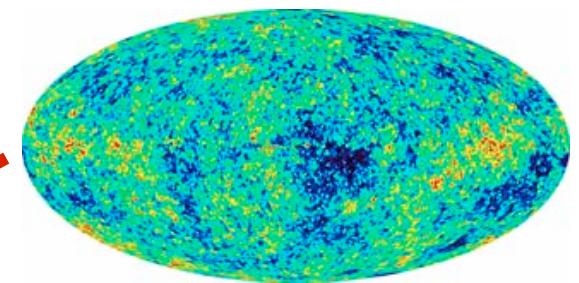


LSS



Our Universe is accelerating

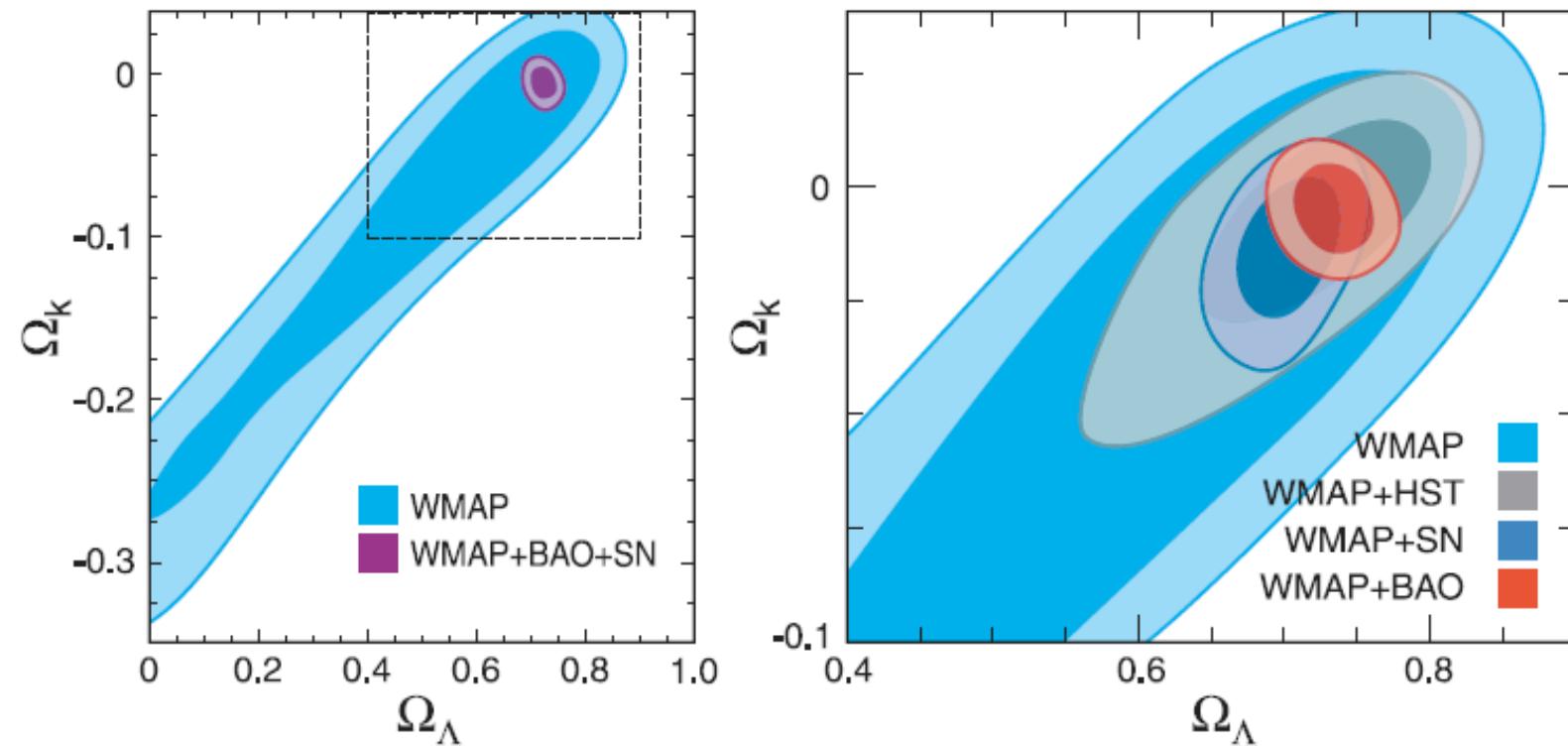
CMB(WMAP)



Universe is mainly dark

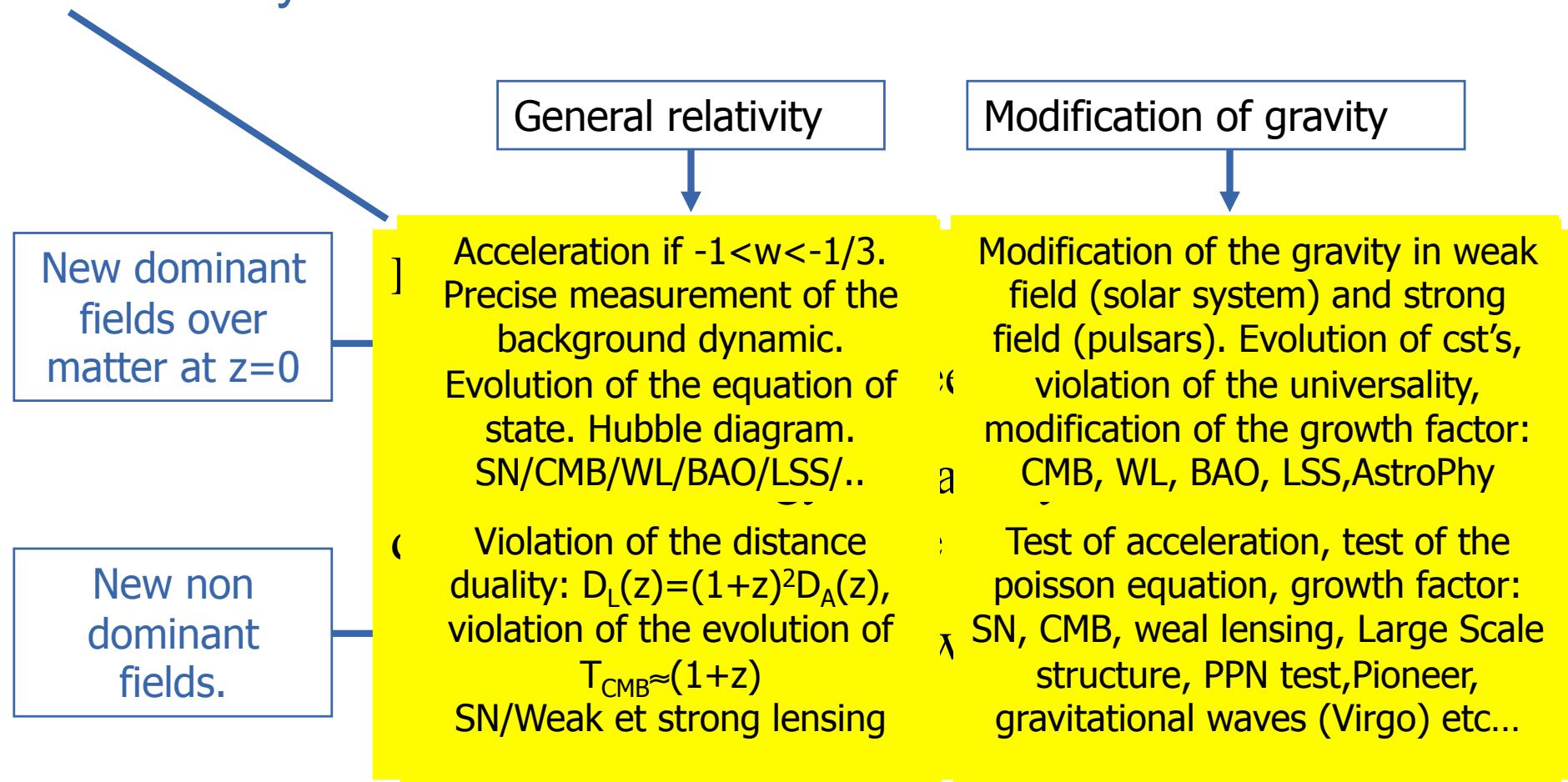
~70% DE  
and  
~25% DM.

a 2006 status



# Phenomenological aspect

- From where this acceleration is coming from ?
- Two main classes of theory/model : Particle physics or Gravity ?

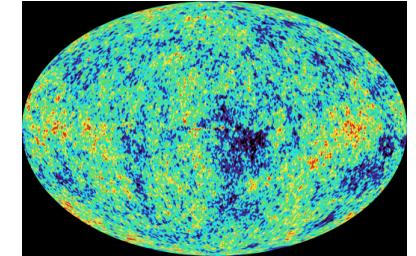


# How to extract $w_0, w_a$ ?

Degeneracies between parameters imply multi-probes analysis sensitive to complementary quantities:

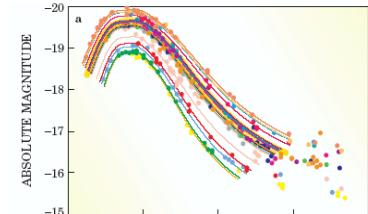
CMB

**Snapshot at ~400,000 yr, viewed from z=0**  
**Angular diameter distance to z~1000**  
**Growth rate of structure (from ISW)**



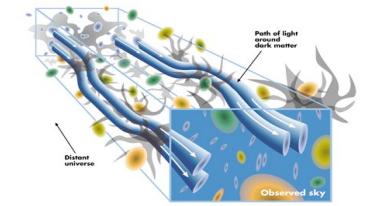
Supernovae

**Standard candle**  
**Luminosity distance**



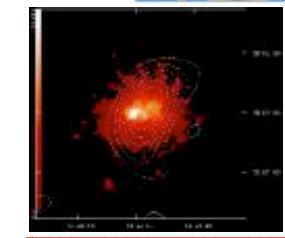
Cosmic Shear

**Evolution of dark matter perturbations**  
**Angular diameter distance**  
**Growth rate of structure**



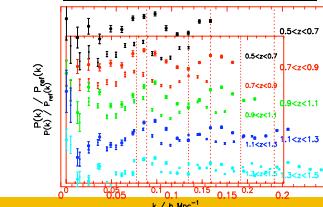
Cluster counts

**Evolution of dark matter perturbations**  
**Angular diameter distance**  
**Growth rate of structure**

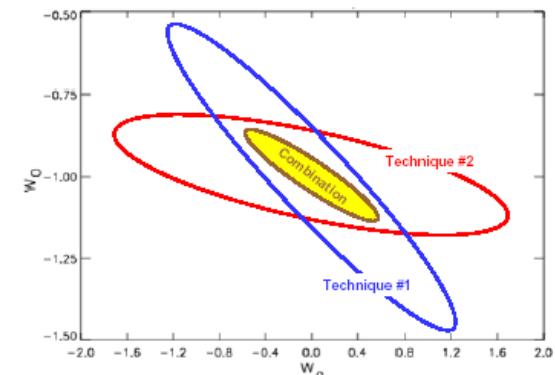


Baryon Wiggles

**Angular diameter distance**



- **Important number of parameters, cosmological and astrophysical**
  - $\Omega_b/\Omega_m$  density for baryon/matter
  - $\Omega_\nu$  density for neutrino's
  - $\Omega_T$  curvature density
  - $H$  Hubble constant,
  - $n_s$  spectral index,
  - $\tau$  reionisation optical depth
  - $\sigma_8$  normalization for CMB, WL and BAO.
  - $m_s$  normalization for SNIa.
  - $y_{he}$  Helium fraction
  - $w_0, w_a$  Equation of state....
  
- **Efficient statistical tools needed:**
  - Bayesian statistic and MCMC : Chinese and astrophysicist choice
  - Frequentist statistic and datagrid : French and particle physicist choice



- Statistic based on  $\chi^2(\Omega_i, w_0, w_a, \dots)$
- Minimum using the gradient method:

$$\nabla \chi^2 / \nabla \Omega_i = 0$$

- Numerical resolution and iterative:

- Error computation:  $U_{kl}^{-1} = \frac{1}{2} \left[ \frac{\partial^2 \chi^2}{\partial \Omega_k \partial \Omega_l} \right]$

- Contour: Solving the equation  $\chi^2 = \chi^2_{\min} + s^2$

Marginalization obtained by minimization:

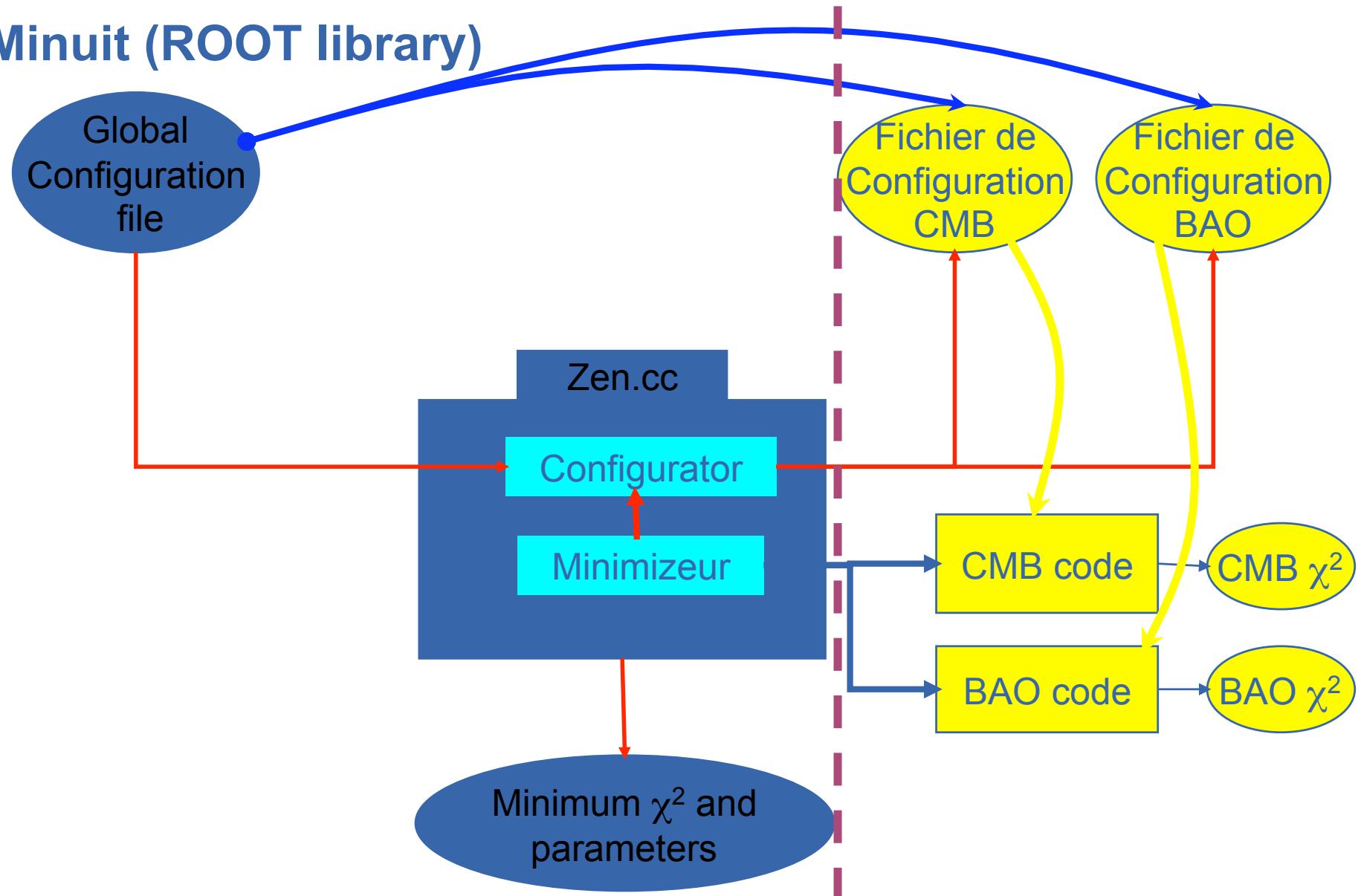
$$\chi^2(w_0, w_a) = \chi^2_{\min}(\Omega_i | w_0, w_a)$$

$$(\Omega_i - \Omega_i^o) = - \left( \frac{\partial^2 \chi^2}{\partial \Omega_k \partial \Omega_i} \Bigg|_{\Omega_k^o} \right)^{-1} \frac{\partial \chi^2}{\partial \Omega_k} \Bigg|_{\Omega_k^o}$$

The contour is constructed by minimizing the  $\chi^2$  on a grid of points (minimum 20\*20) and iso- $\chi^2$  are constructed using interpolation. Each point (20 hours of computing) is calculated on a CE.

**A simple contour requires about 1 year of CPU on a single CPU.  
Thanks to datagrid (result in few days)**

## Minuit (ROOT library)



# Global configuration file

```

action = compute_chi2/find_minimum
simulation = no
#-----
scan  = wa
from  = 0.
upto  = 0.1
nbscan = 100
# contour ?
scan2d = none
from2d = -1.0
upto2d = 0.0
nbscan2d = 2
#-----
use_sn = yes
use_bao = yes
use_wl = no
use_cmb = yes
use_hst = no
cmb_code = cmbeasy
flatness =yes
#-----
```

```

#-----
fit_omega_b = yes
fit_omega_cdm = yes
fit_omega_nu = no
fit_ns = yes
fit_w0 = yes
fit_wa = yes
.....
#-----
perturb_type = nopersp
gauge = quintsynchronous
.....
# Equation of state
w0      = -1.095936698
min_w0  = -3.
max_w0  = -0.3
error_w0 = 0.2
# Equation of state variation.
wa      = -0.2295558424
min_wa  = -2.5
max_wa  = 2.5
error_wa = 0.8
```

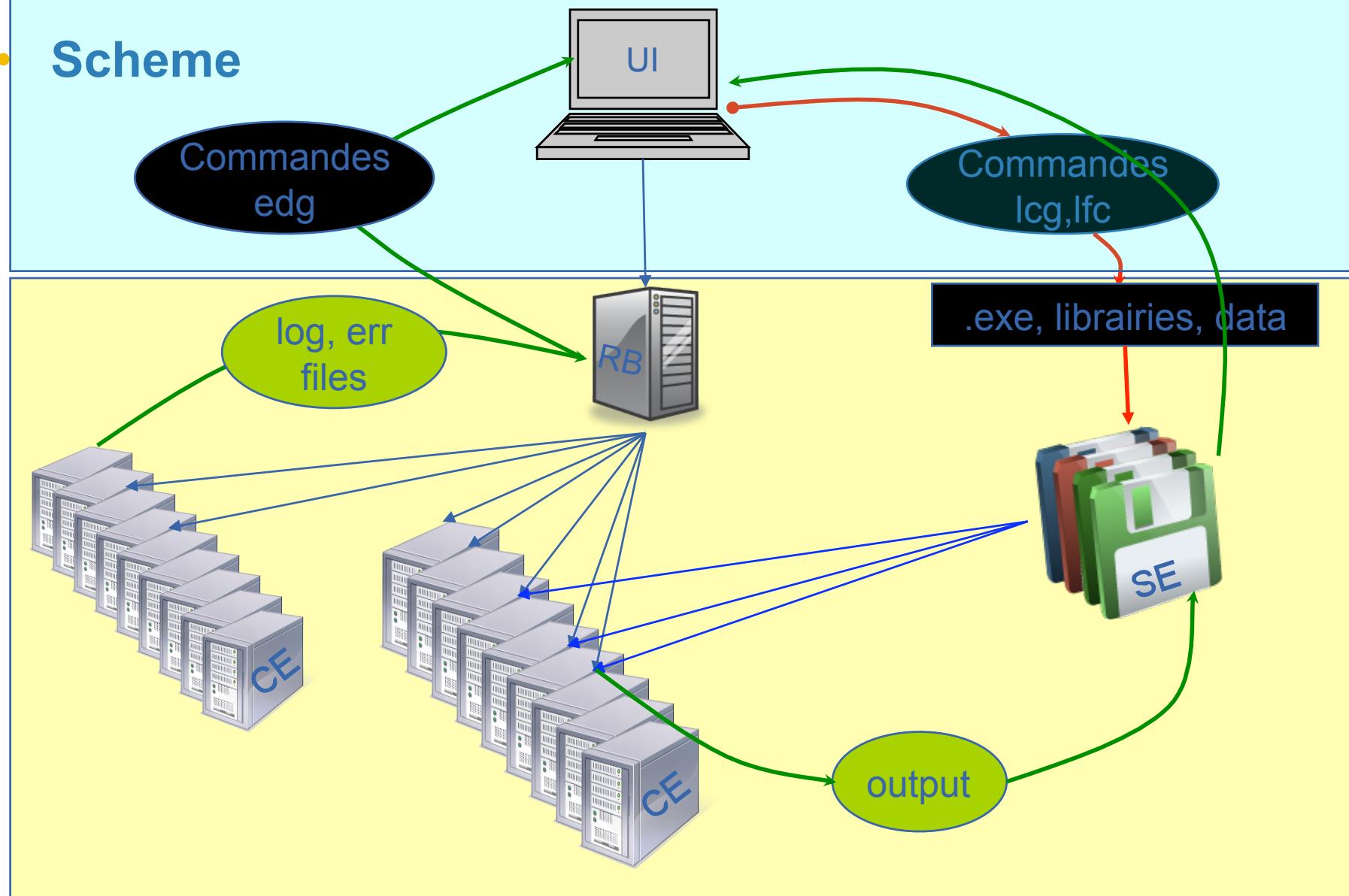
- Format totalement libre (fichier texte)
- Héritage des paramètres globaux

```
# General configuration file
.....
#-----
perturb_type = noperturbsup
gauge = quintsynchronous
.....
# Equation of state
w0      = -1.095936698
min_w0   = -3.
max_w0   = -0.3
error_w0 = 0.2
# Equation of state variation.
wa       = -0.2295558424
.....
```

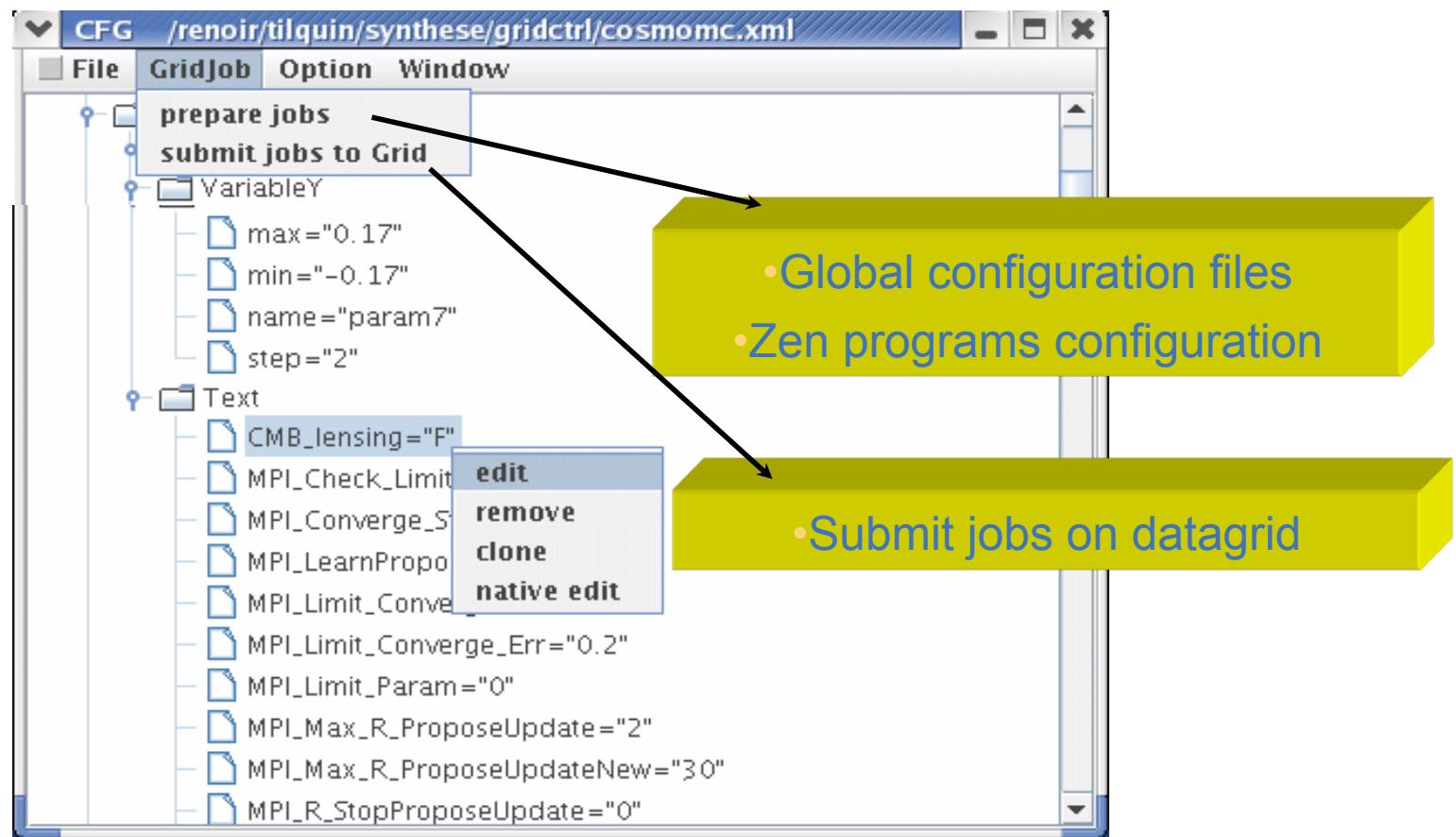
configuteur

```
# Configuration file for BAO
omegab = [omega_b]
omegam = [omegam]
h = [h]
ns = [ns]
tau = [tau]
w0 = [w0]
wa = [wa]
sigma8 = [sigma8]
bias = [bao_bias]
#-----
Internal_flag = 1
#     Should be zero. Curvature not
#     yet implemented in bao
omegak = 0
.....
```

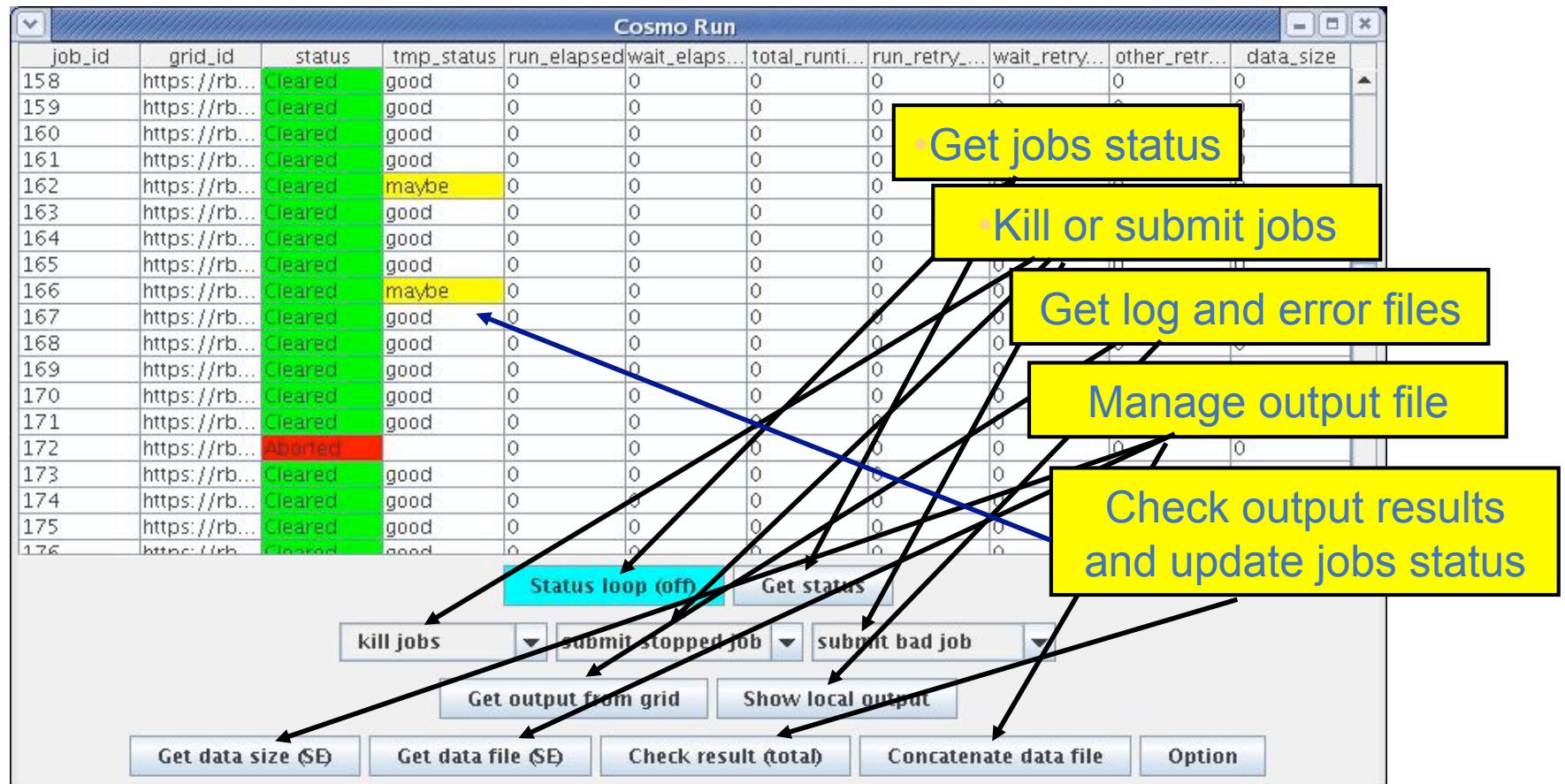
- Scheme



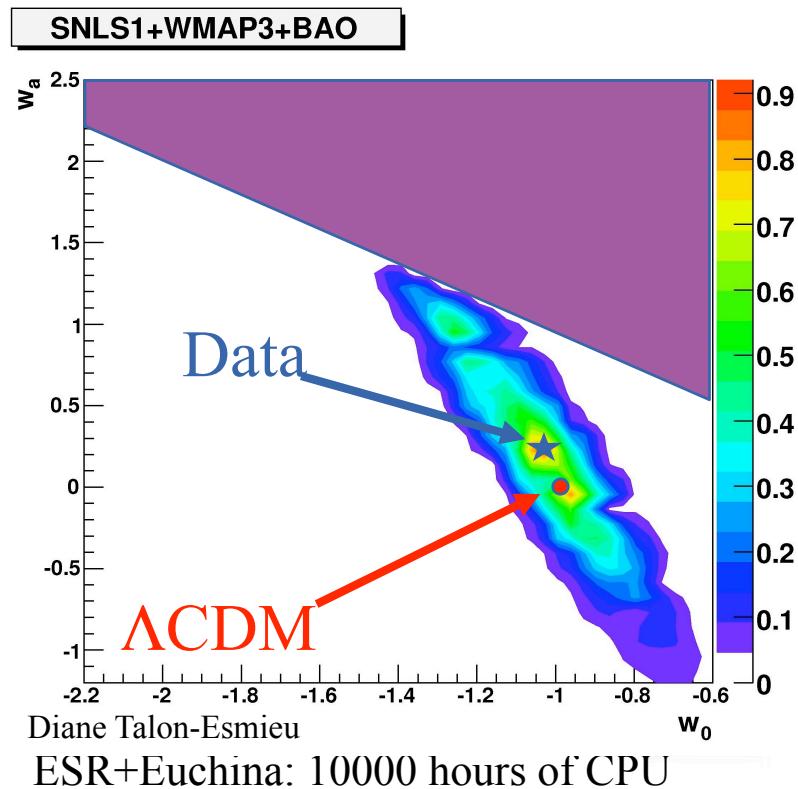
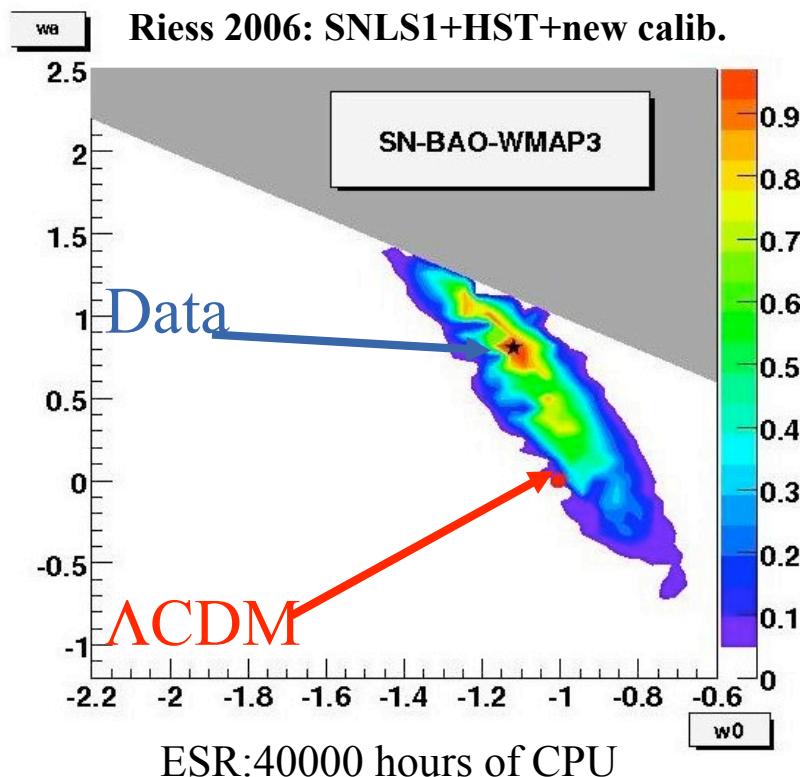
- A run is a set of  $n \times n$  jobs (400) send in parallel. A graphical interface has been developed (thanks to Zuxuan Qian) to construct and submit them.



- Jobs monitoring and output data manager

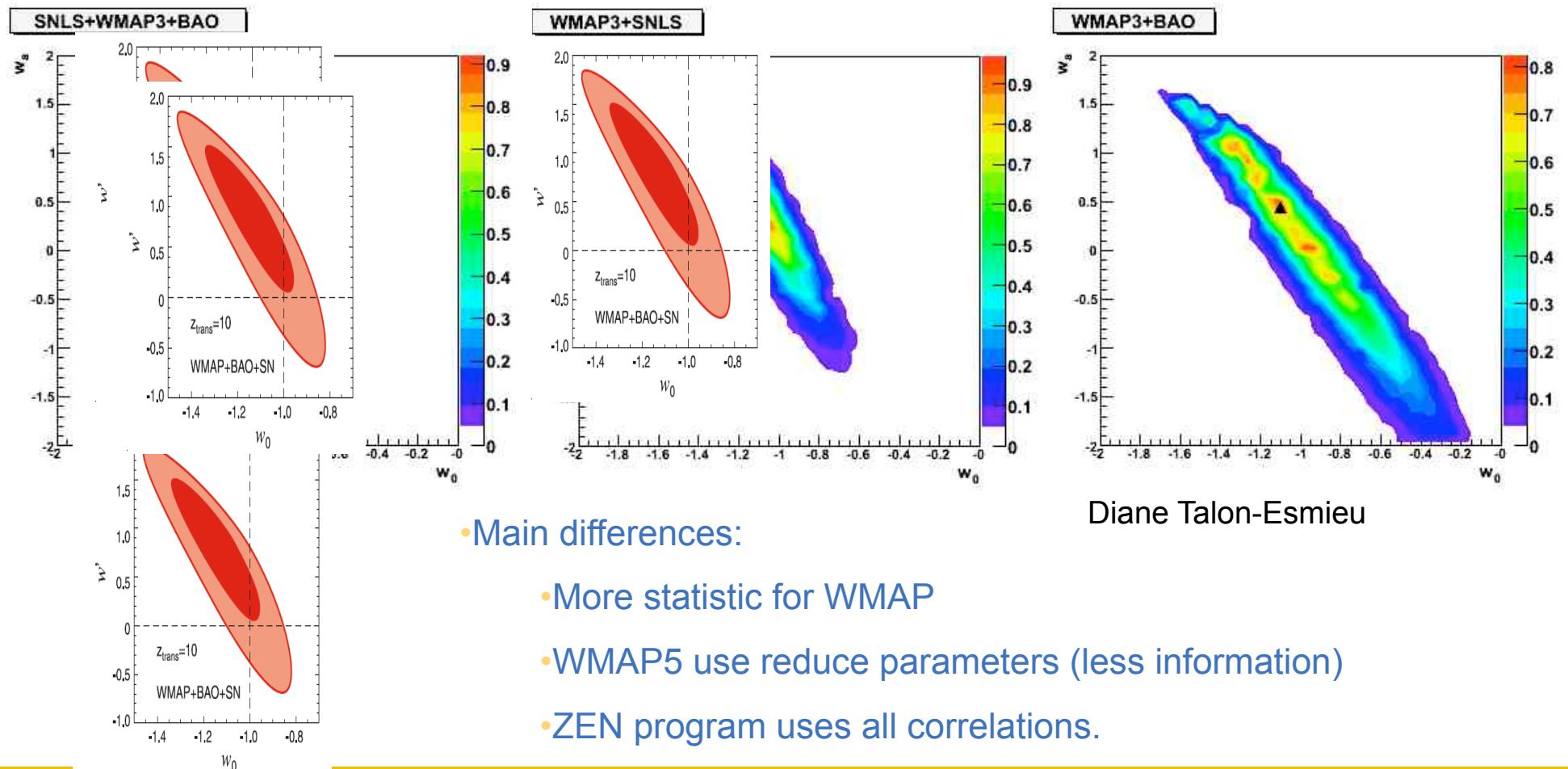


- First results using datagrid within:
  - ESR (Earth Science Research) VO (thanks to M. Petit-Didier) and Euchina Virtual organization
  - Using SN+CMB+BAO with frequentist statistic and (BAOfit from Sun Lei, PKU-CPPM Join PhD)



- News results on WMAP3-SNLS-BAO comparison

Comparison with WMAP5+SNLS+BAO last publication



- The ZEN program is now running on datagrid  
Thanks to ESR virtual organization
- The two graphical interfaces are very powerfull tools to submit jobs and to debug.
- Major problem was the linux version (SL3,SL4)
- Our first results on multi-probes analysis are compatible with competitive analysis (Bayesian) and WMAP5
- Perspectives:
  - Add new probes (WL, etc...) and add new data (WMAP5,SNLS3...)
  - Investigate new theoretical models and implement them in ZEN
  - Add new statistics (MCMC, Bayesian...)
  - Start systematic analysis !
  - Write a documentation : IT IS A PUBLIC TOOL
    - If you want to use it you are welcome (possible training).